

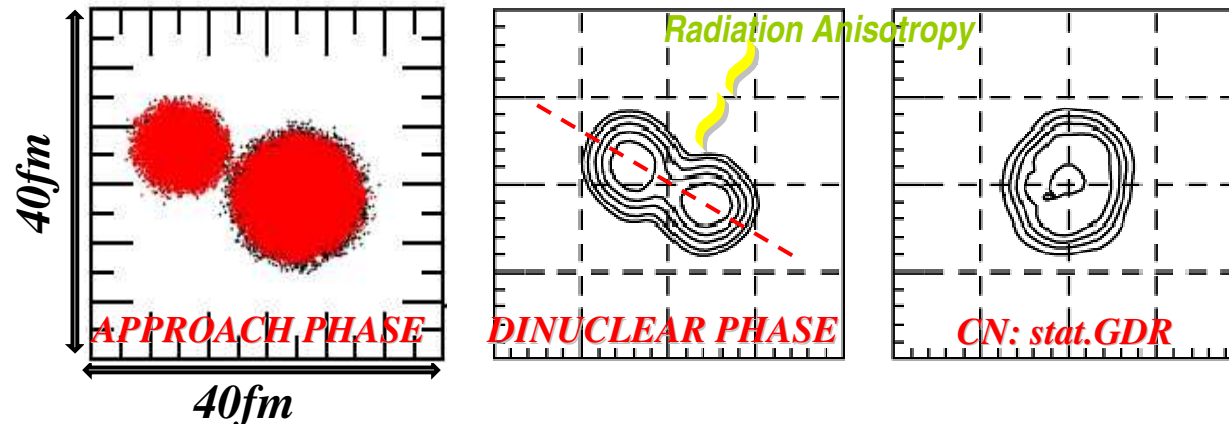
***“PRE-EQUILIBRIUM DIPOLE
RADIATION WITH PROTON-RICH
BEAMS”***



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LNS - INFN

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CHARGE EQUILIBRATION DYNAMICS:

• **STOCHASTIC** → Diffusion.

VS.

• **COLLECTIVE** → Dipole Oscillations of the Di-Nuclear System in Fusion Dynamics between colliding ions with different N/Z ratios.



DYNAMICAL DIPOLE: Pre-equilibrium dipole mode in phase dinuclear give rise to a prompt γ -ray emission which depends:

• on initial dipole moment,
$$D(t_0) = \frac{Z_1 Z_2}{A} \left(\frac{N_1}{Z_1} - \frac{N_2}{Z_2} \right) (R_1 + R_2);$$

• on the symmetry term ($< \rho_0$), **Isvector Restoring Force** → Centroid, Yield
Neutron emission, NN cross sections → Damping

• on the fusion dynamics, **Neck Dynamics (Mass Asymmetry)**

Anisotropy

Cooling on the way to Fusion

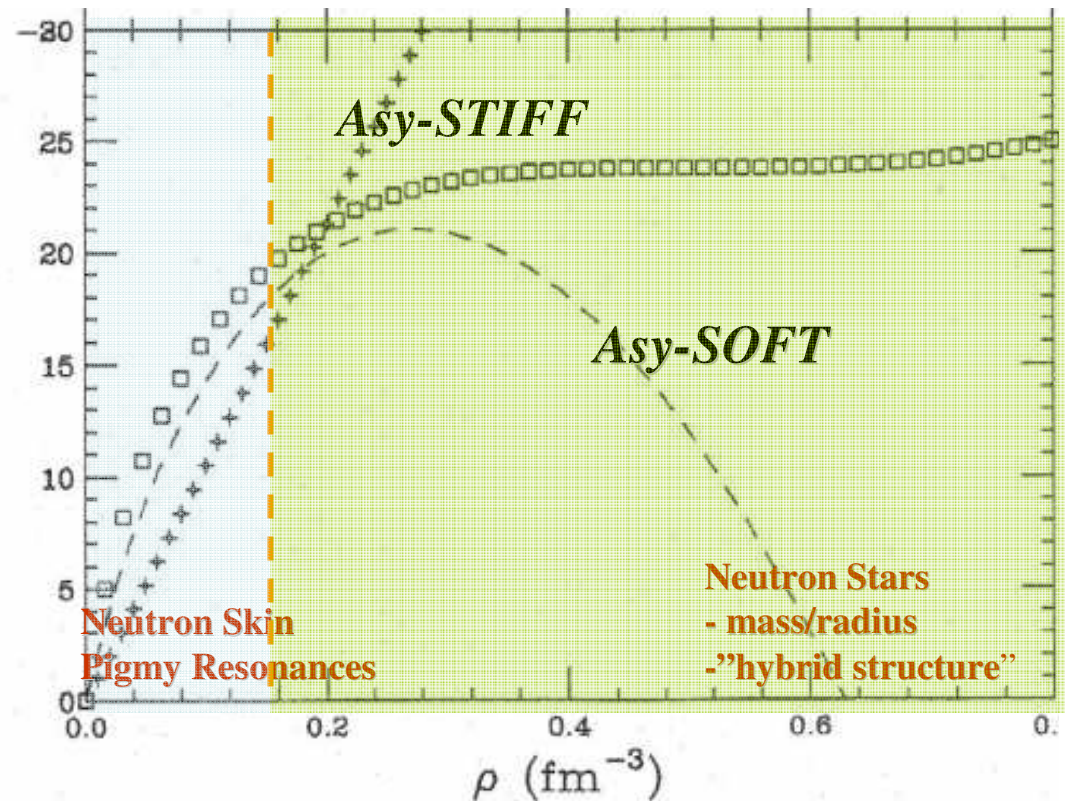
Reactions have been simulated by considering a stochastic extension of microscopic transport equation BNV, following a test-particle evolution:

•VLASOV + NN-COLLISIONS and PAULI CORRELATIONS

$$\frac{E}{A}(\rho, I) = \frac{E}{A}(\rho, 0) + \frac{E_{\text{sym}}}{A}(\rho) I^2$$

$$\frac{E_{\text{sym}}}{A}(\rho) = \frac{E_{\text{sym}}^{\text{KIN}}}{A}(\rho) + \frac{E_{\text{sym}}^{\text{POT}}}{A}(\rho)$$

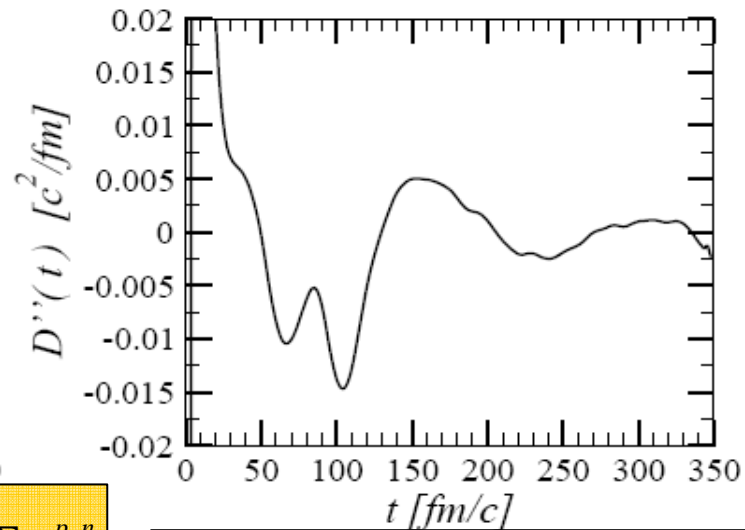
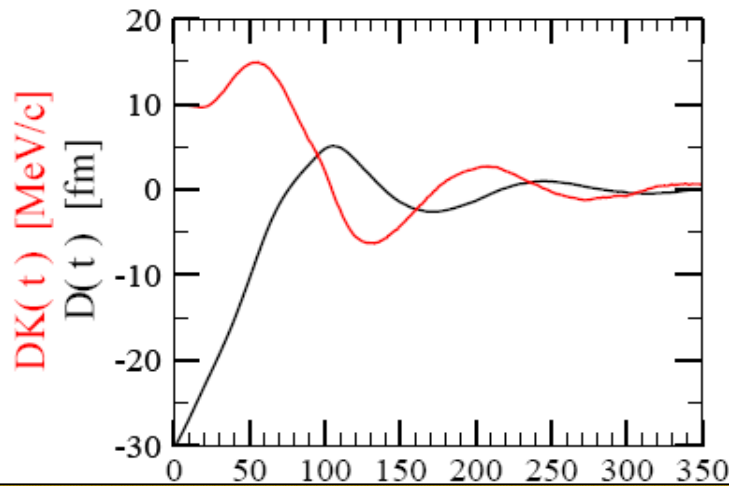
$E_{\text{pot}}^{\text{sym}} / A \text{ (MeV)}$



- The isoscalar section of the EoS is the same in both cases, corresponding to a compressibility around 220MeV.
- Numerical approach with 200 gaussian test particles per nucleon.
- In the collision integral in medium N-N cross sections are considered.
- To reduce the numerical noise we run fifty events for each set of macroscopic initial conditions and the displayed quantities are the averages over this ensemble

Pre-equilibrium dipole emission → Bremsstrahlung - Quantitative estimations

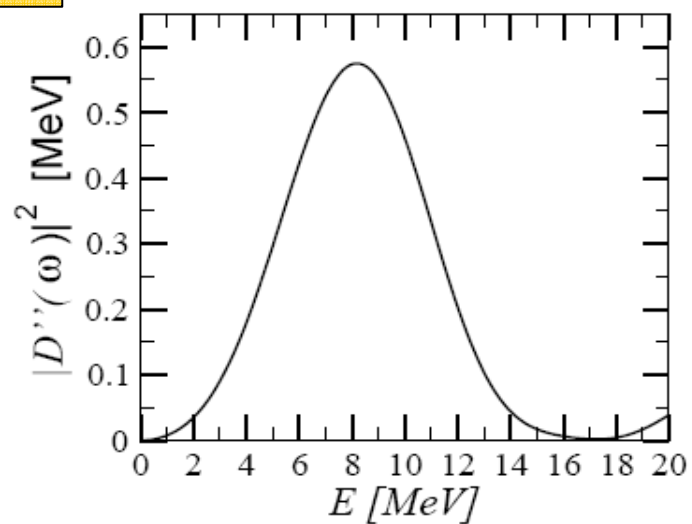
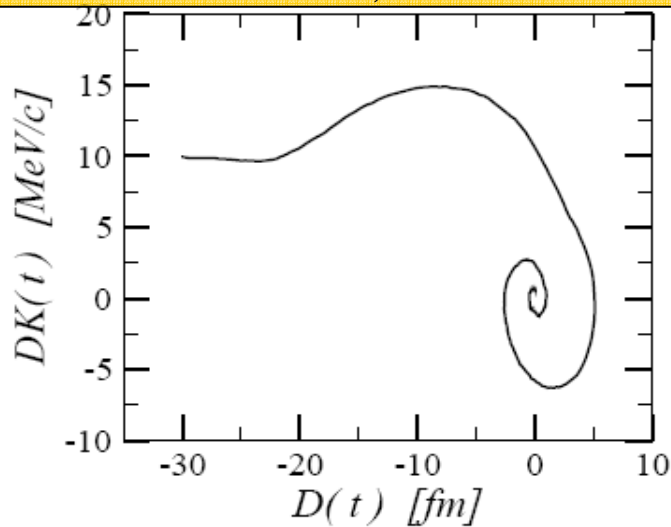
$^{36}\text{Ar} + ^{96}\text{Zr}$ 16A MeV $b = 4\text{fm}$ Asystiff



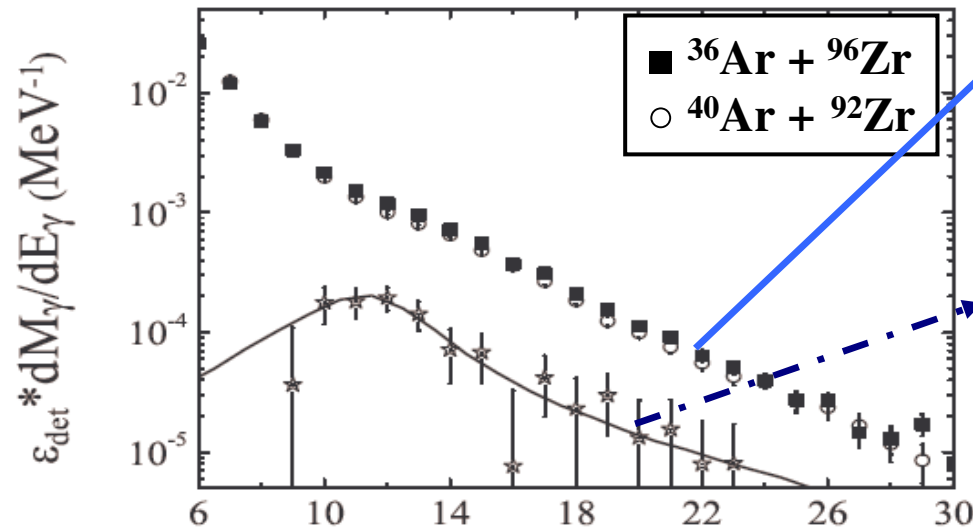
$$D(t) \equiv \frac{NZ}{A} [X_p(t) - X_n(t)] \xrightarrow{[fm/c]} X_{p,n} \equiv \frac{1}{Z, N} \sum x_i^{p,n}$$

$$DK(t) \equiv P_p - P_n \rightarrow P_{p,n} \equiv \frac{1}{Z, N} \sum p_i^{p,n}$$

$$\frac{dP}{dE_\gamma} = \frac{2e^2}{3\pi\hbar c^3 E_\gamma} \left(\frac{NZ}{A} \right)^2 |D''(\omega)|^2$$

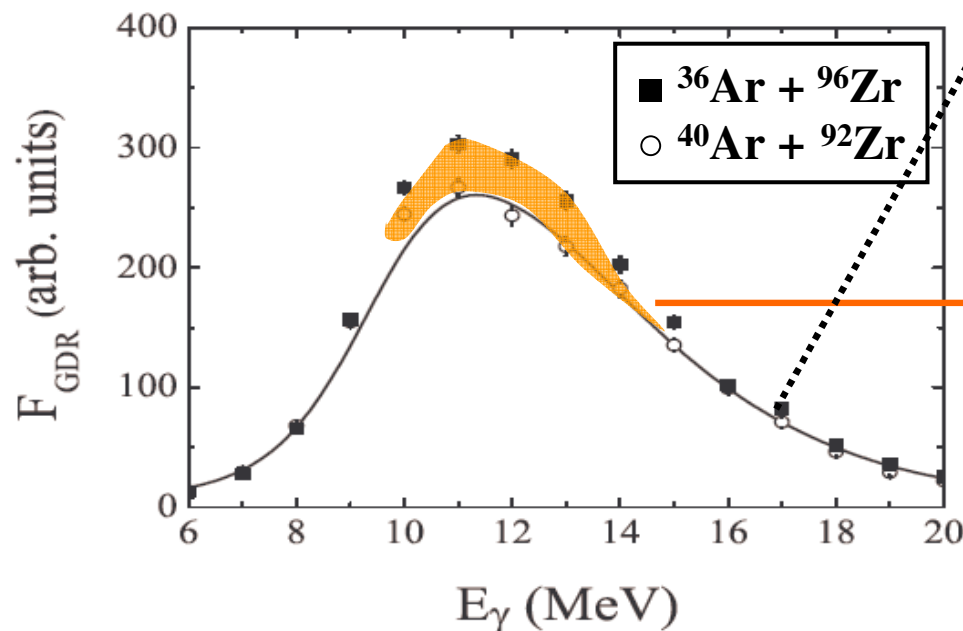


$^{36}\text{Ar}+^{96}\text{Zr}$ vs. $^{40}\text{Ar}+^{92}\text{Zr}$ at 16 MeV beam energy. Fusion events: same CN selection



• Bremsstrahlung-subtracted γ -ray spectra at $\theta_\gamma=90^\circ$ (vs. Beam Axis) in coincidence with the evaporation residues.

• Difference.



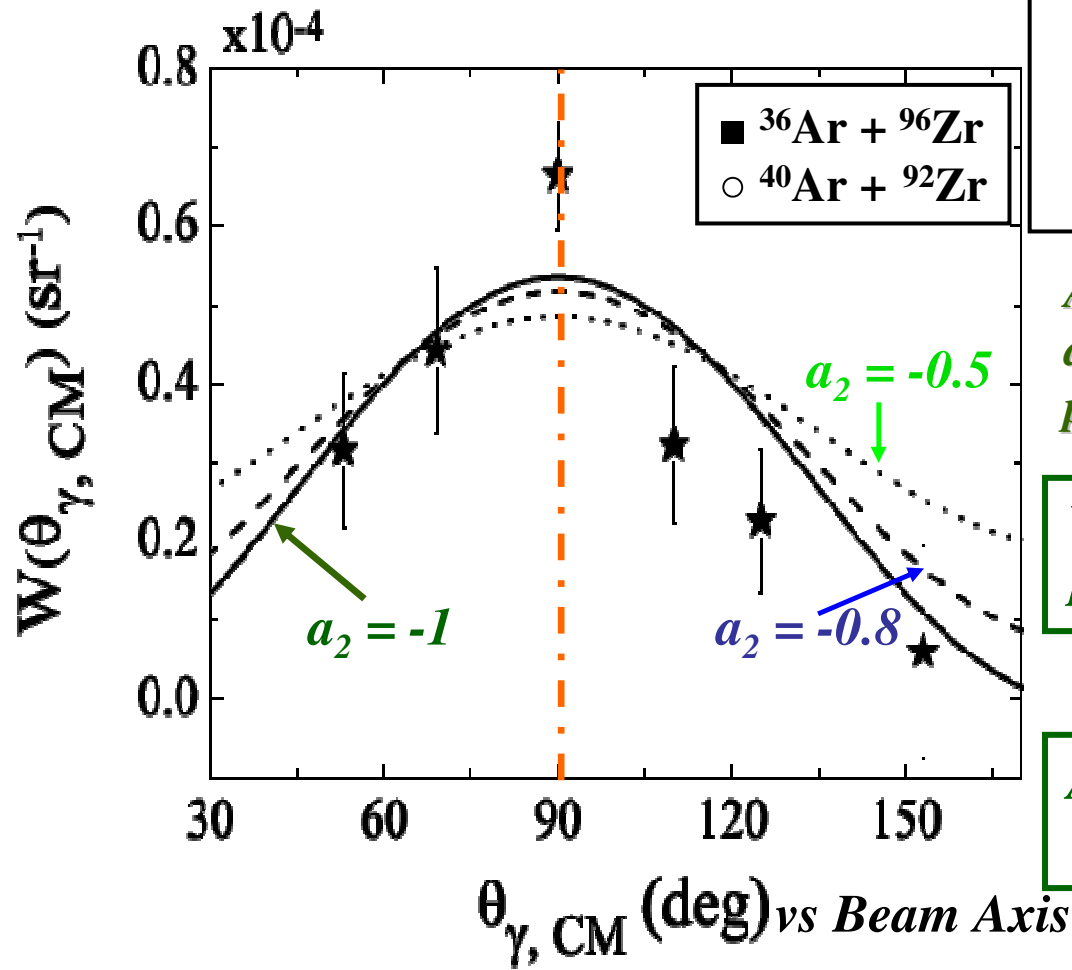
The data are linearized by theoretical γ -ray spectrum of the charge symmetric reaction calculated with the statistical decay code CASCADE results without a CN GDR

An increase of the GDR γ -ray intensity is found by going toward the more charge asymmetric system. θ_γ - study of the extra-yield with MEDEA

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Gamma Angular Distribution, $10 < E_\gamma < 15$ MeV

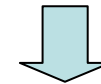
$$W(\vartheta_\gamma) = W_0 [1 + a_2 P_2(\cos \vartheta_\gamma)]$$



$W(\Theta_{\gamma, \text{CM}})$ is anisotropic with a maximum around 90° and it is consistent with emission from a dipole oscillating along the beam axis.

A strong dipole-like photon angular distribution has been observed, with the parameter $a_2 = -1$, $W(\Theta_{\gamma, \text{CM}}) \sim \sin^2 \Theta_{\gamma, \text{CM}}$

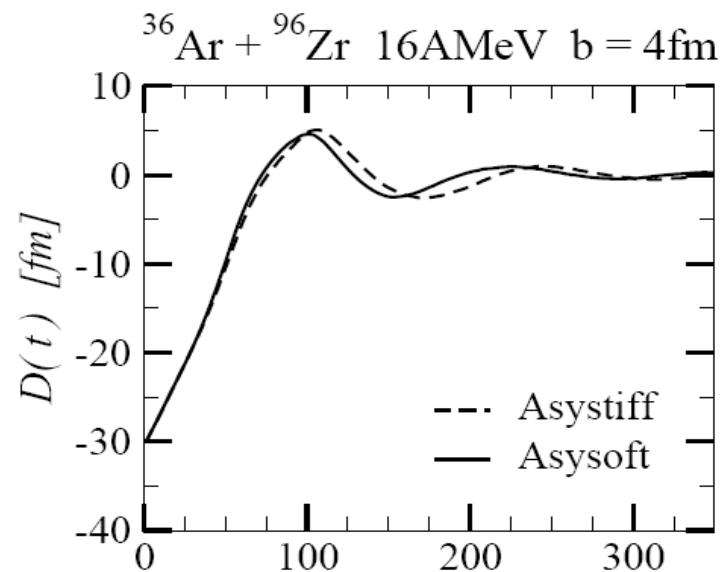
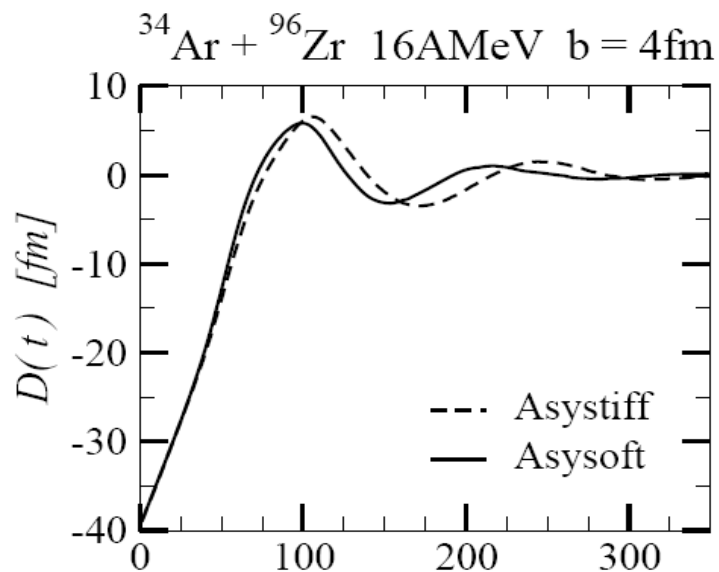
Widening: rotation of the Prompt Dipole Axis vs the Beam Axis



Accurate Angular Distrib. Measure: Dipole Clock!

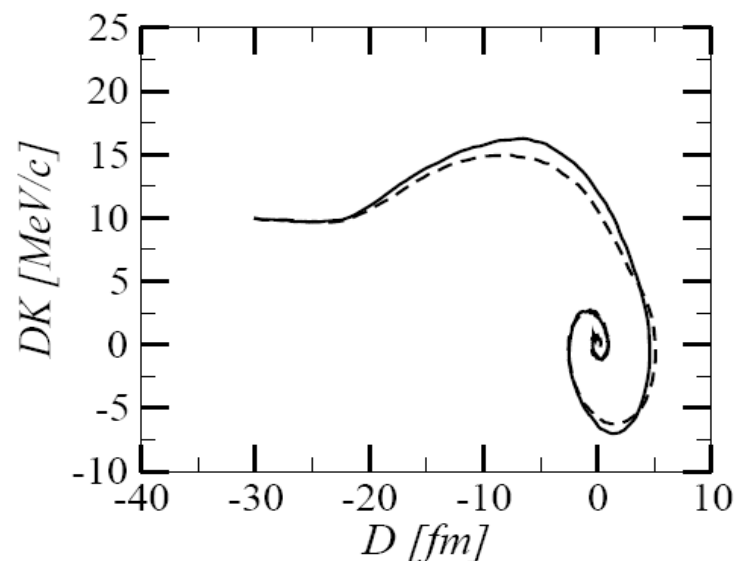
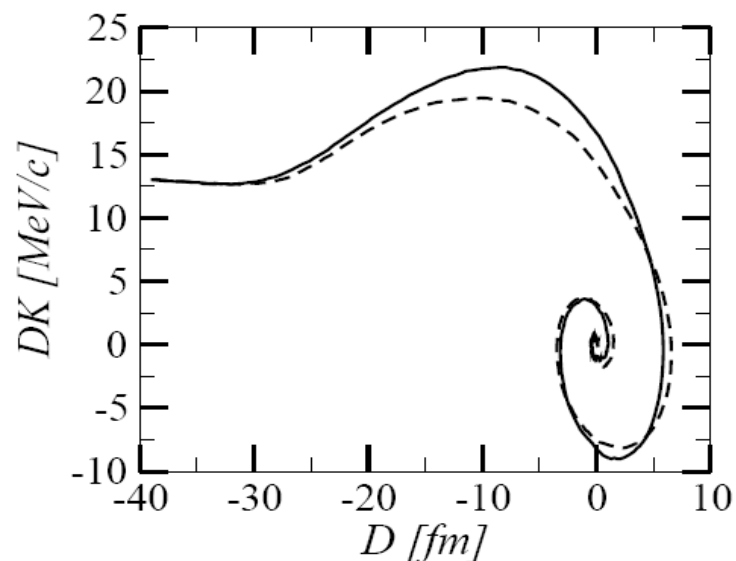
We study the features of the preequilibrium dipole considering the reaction $^{36}\text{Ar}+^{96}\text{Zr}$ at 16A MeV , as referred to the same reaction induced by a ***p-rich ^{34}Ar beam***.

Prompt Dipole Oscillations

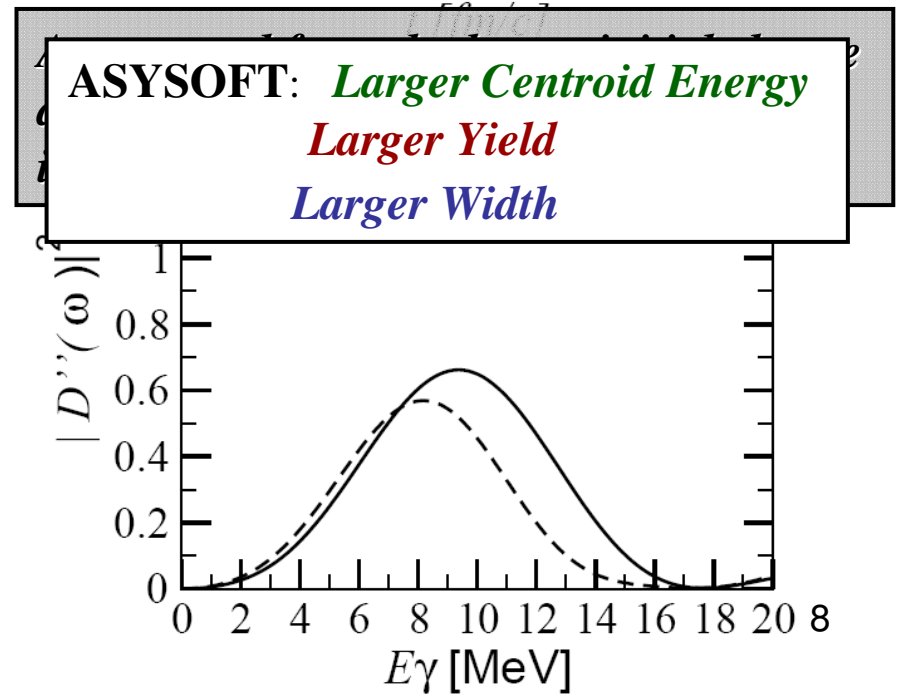
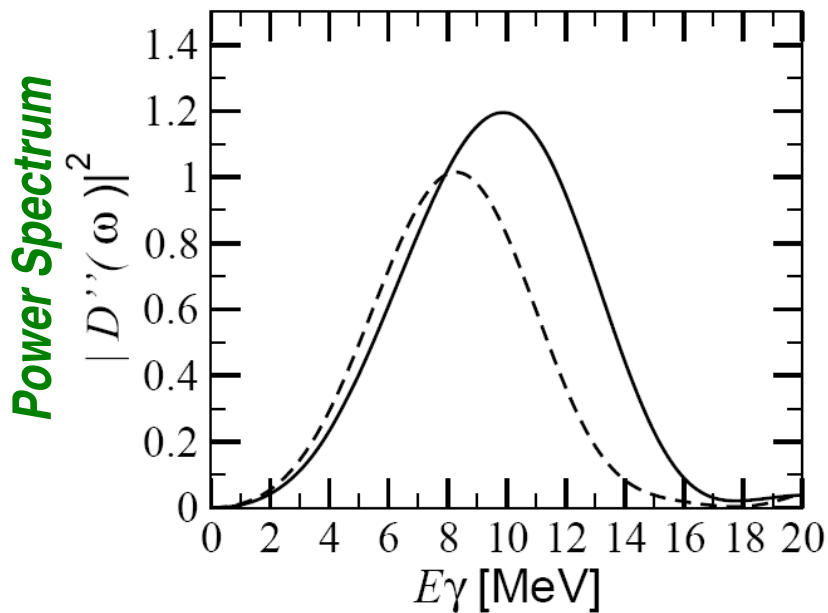
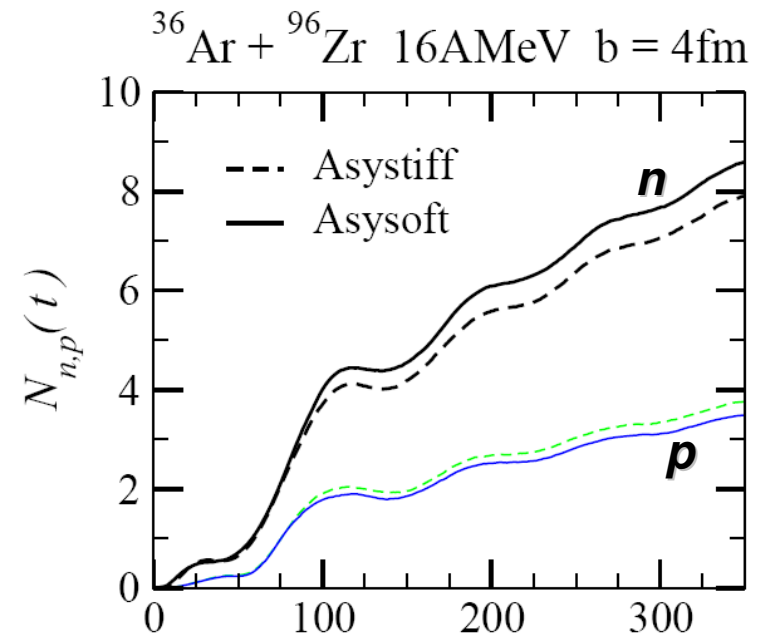
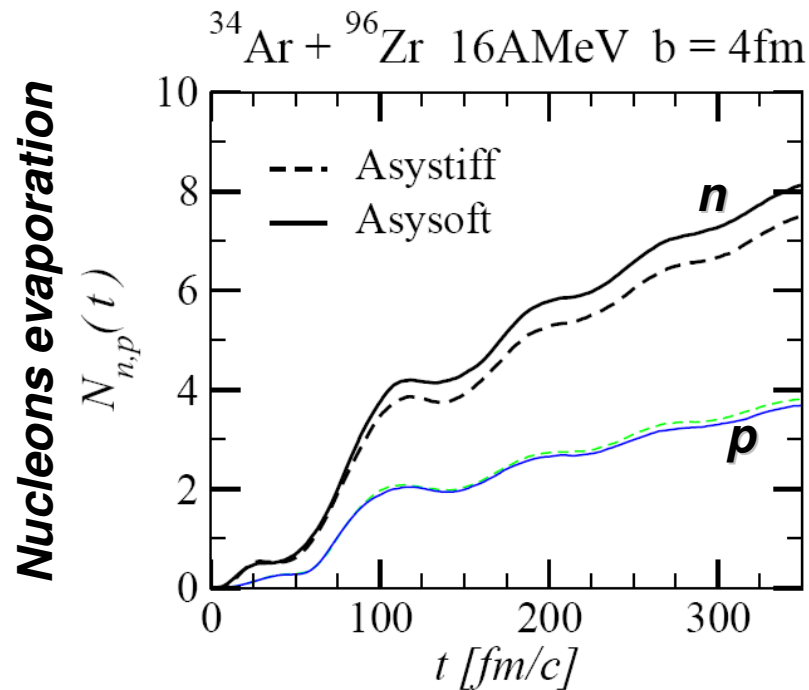


$D(t_o)$ has a value around **27fm** for the **^{34}Ar** beam, and a value **20fm** for the **^{36}Ar** beam

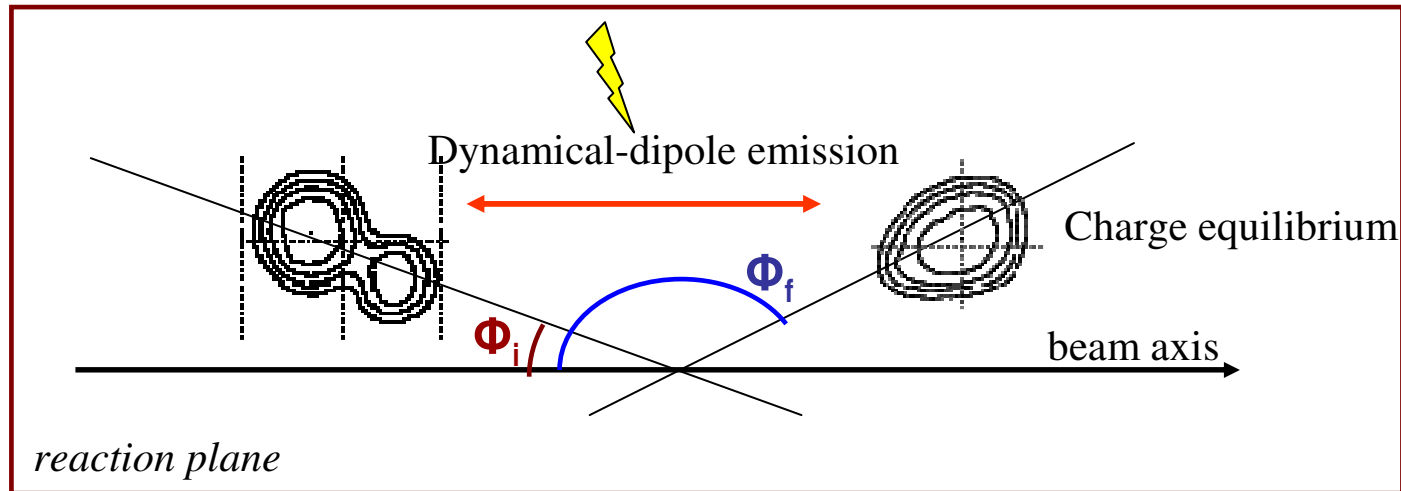
Phase Space Correlations



ASYSOFT: Neutrons see a more repulsive Symmetry Potential below ρ_0



Angular distribution: Rotation on the Reaction Plane of the Emitting Dinuclear System



$$W(\vartheta_\gamma) = W_0 \left[1 + a_2 P_2(\cos \vartheta_\gamma) \right], \quad a_2 = -\left(\frac{1}{4} + \frac{3}{4} x \right) \quad \theta_\gamma : \text{photon angle vs. beam axis}$$

$$x = \cos(\Phi_i + \Phi_f) \frac{\sin(\Delta\Phi)}{\Delta\Phi}, \quad \Delta\Phi = \Phi_f - \Phi_i$$

• All probed Rotating angles: $\Delta\Phi = 2\pi \rightarrow x = 0 \rightarrow a_2 = -1/4$

➡ Statistical result, Collective Prolate on the Reaction Plane

• Fixed Rotating angle: $\Delta\Phi = 0 \rightarrow \Phi_i = \Phi_f = \Phi_0 \rightarrow W(\vartheta_\gamma) \propto [1 - (1 - \sin^2 \Phi_0) P_2(\cos \vartheta_\gamma)]$

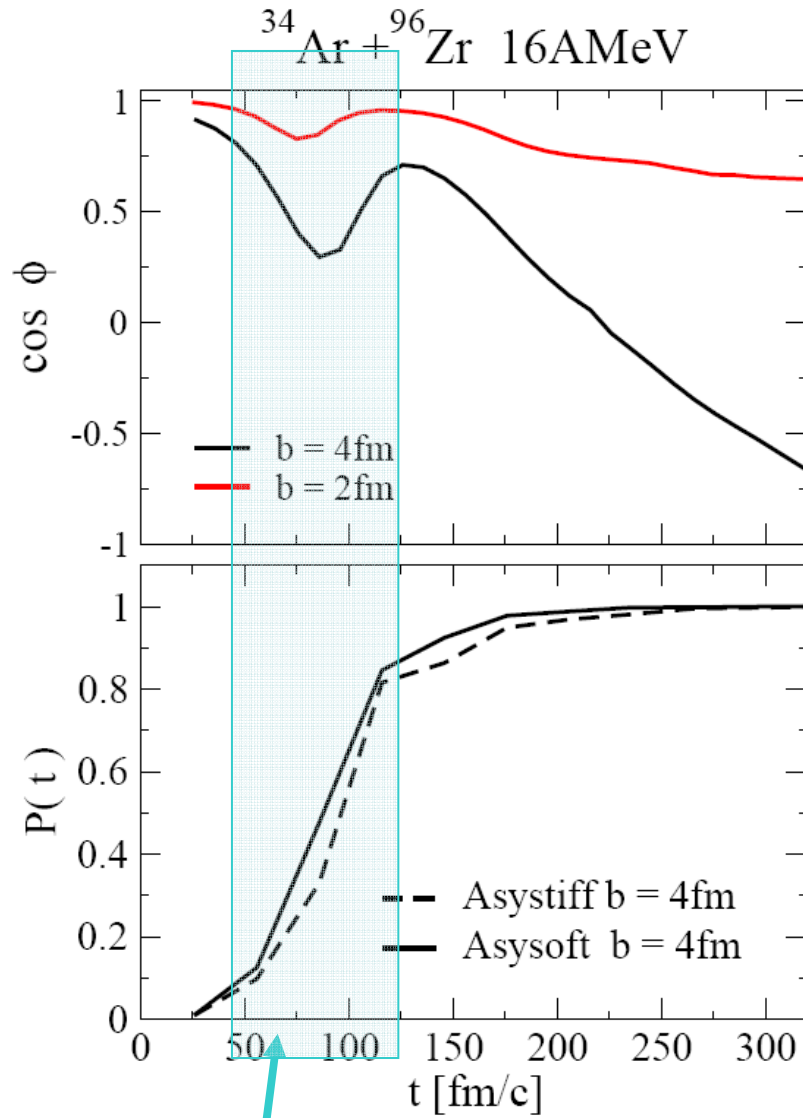


• No rotation: $\Phi_0 = 0 \rightarrow \sin^2 \theta_\gamma$
Pure Dipole



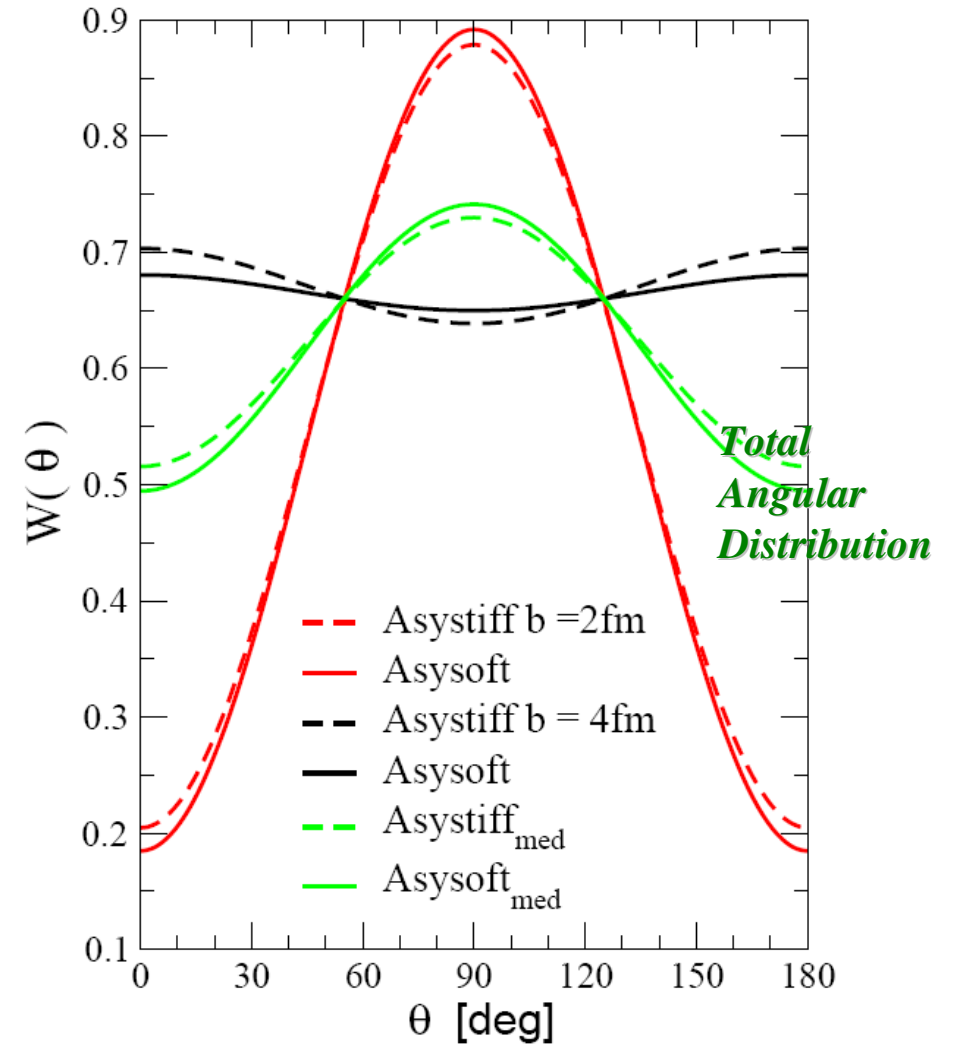
Dominant since the prompt dipole is rapidly damped? Angular Distribution
→ Dynamic Dipole Lifetime

Rotation Angle



Emission Probability: delayed with asystiff

$$P(t) = \int_{t_0}^t |D''(t)|^2 dt / P_{tot}$$

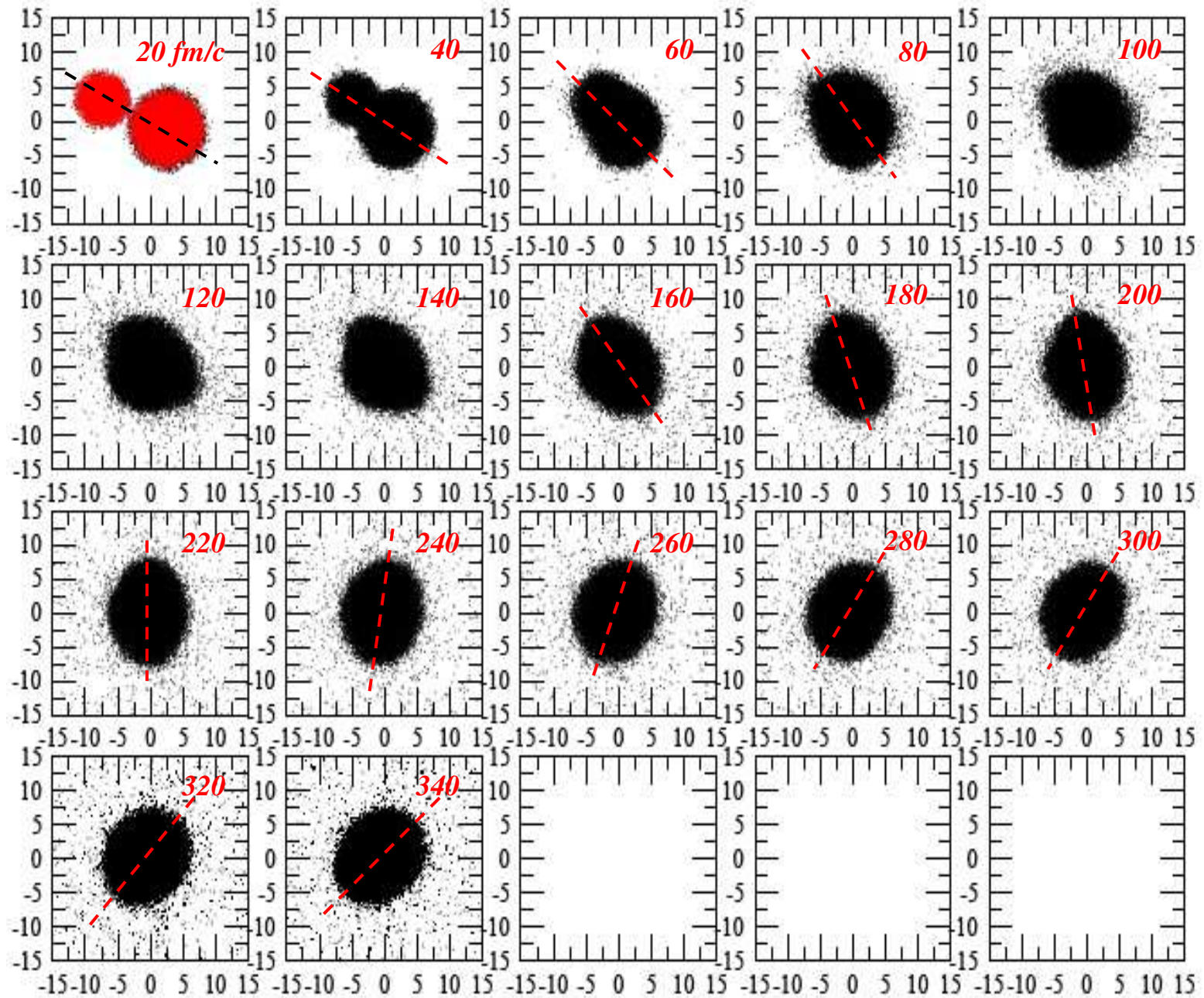


Weighted anisotropies:

$$W(\theta) = \sum_{i=1}^{t_{max}} \beta_i W(\theta, \Phi_i)$$

$$\beta_i = P(t_i) - P(t_{i-1})$$

$^{34}\text{Ar} + ^{96}\text{Zr}$ 16Amev $b = 4\text{fm}$ Asystiff



Conclusions

The dynamical dipole mode appears to be a suitable probe to test the symmetry energy term in the nuclear EoS as well as to scrutinize the early entrance channel dynamics in dissipative reactions with p-rich beams

The angular distributions are also sensitive to the fusion dynamics and dipole excitation mechanism and lifetime.

We can get the total yield for the two Iso-EoS and the two systems :

- We obtain 1.22 for ^{34}Ar reaction and 1.17 for ^{36}Ar in the Asysoft/Asystiff case;*
- We obtain a value around 1.7 for Asystiff choice and 1.8 for Asysoft in the $^{34}\text{Ar}/^{36}\text{Ar}$ reaction case.*
- Use of a more symmetric reference system $^{40}\text{Ar} + ^{90}\text{Zr}$, $D(t_0)=1.6\text{fm}$.*

Beam intensity problem:

- exp. Medea $^{36}\text{Ar} + ^{96}\text{Zr}$ intensity $2\text{nA} \rightarrow \sim 10^9\text{pps}$;*
- new exp with $^{34}\text{Ar} \rightarrow 10^6\text{-}10^8 \text{ ??}$*